



Leibniz-Institut für
Astrophysik Potsdam

Overview and preliminary results of a near-infrared cross-dispersed spectrograph based on the arrayed waveguide grating technology

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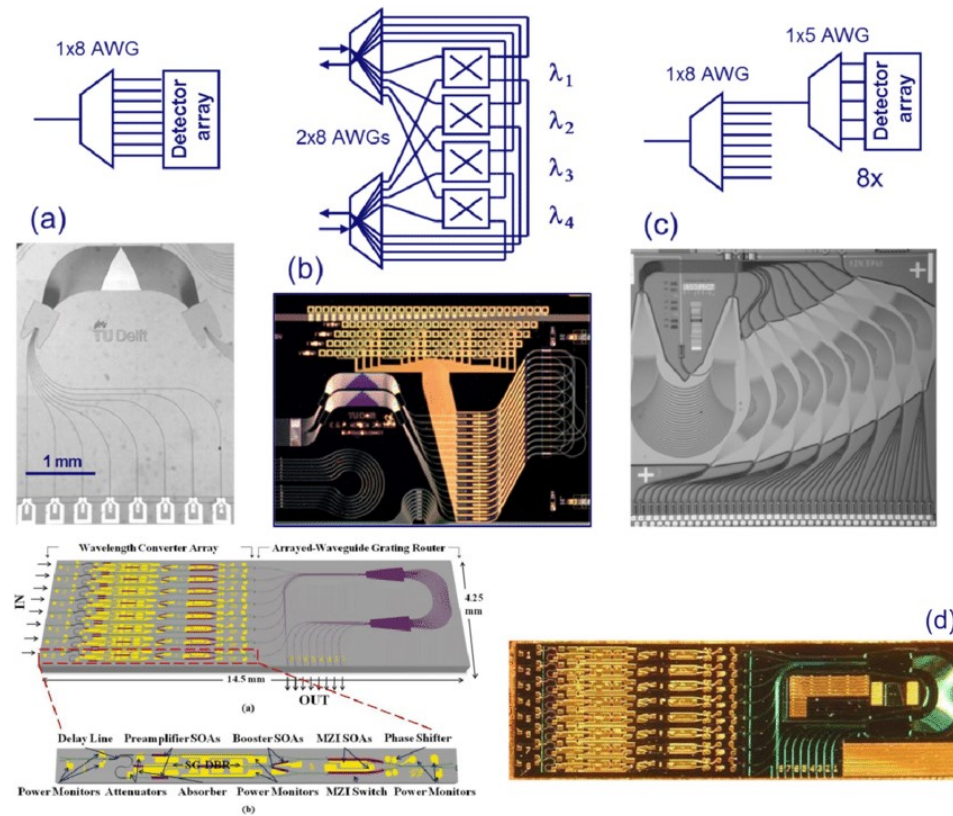
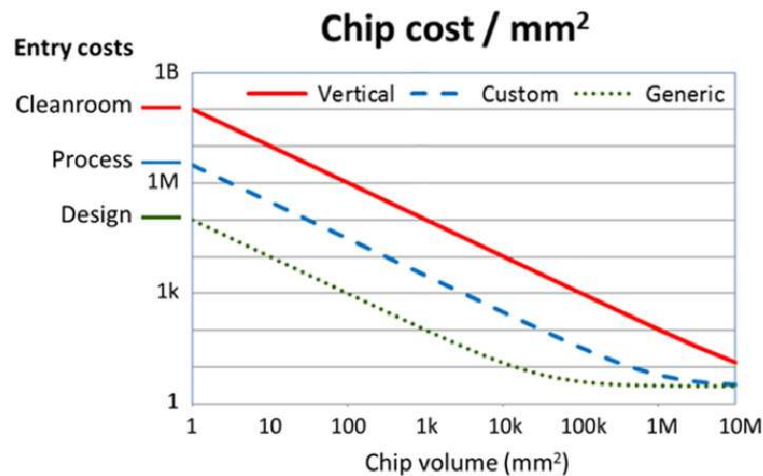
09.06.2023

Outline

- 1) PICs Overview
- 2) Applications in Astronomy
- 3) Astrophotonics at AIP – Arrayed Waveguide Gratings
- 4) The Potsdam Arrayed Waveguide Spectrograph (PAWS)
- 5) POCO Frequency Comb System – PAWS Calibration
- 6) Outlook

PICs Overview

- Development of chip complexity over the years
 - Similar to Moors-Law in transistors on microchips
- PICs are widely used
 - WDM receivers, Mach-Zehnder, etc...
- Dependence of chip cost
 - Chip cost on the production for three models



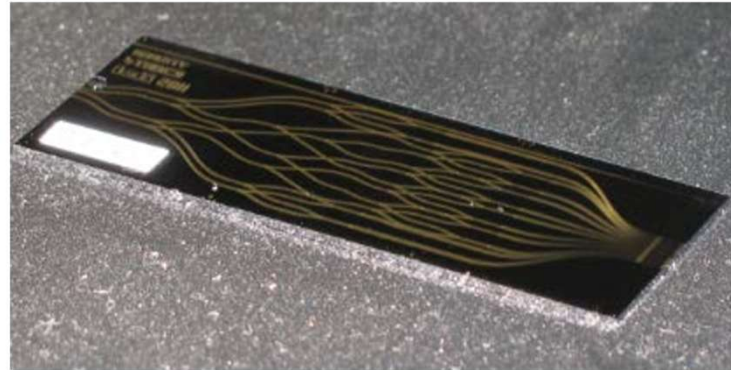
Example of different PICs

PICs in Astronomy

- PICs benefits



Beam combination at the AMBER instrument,
VLT

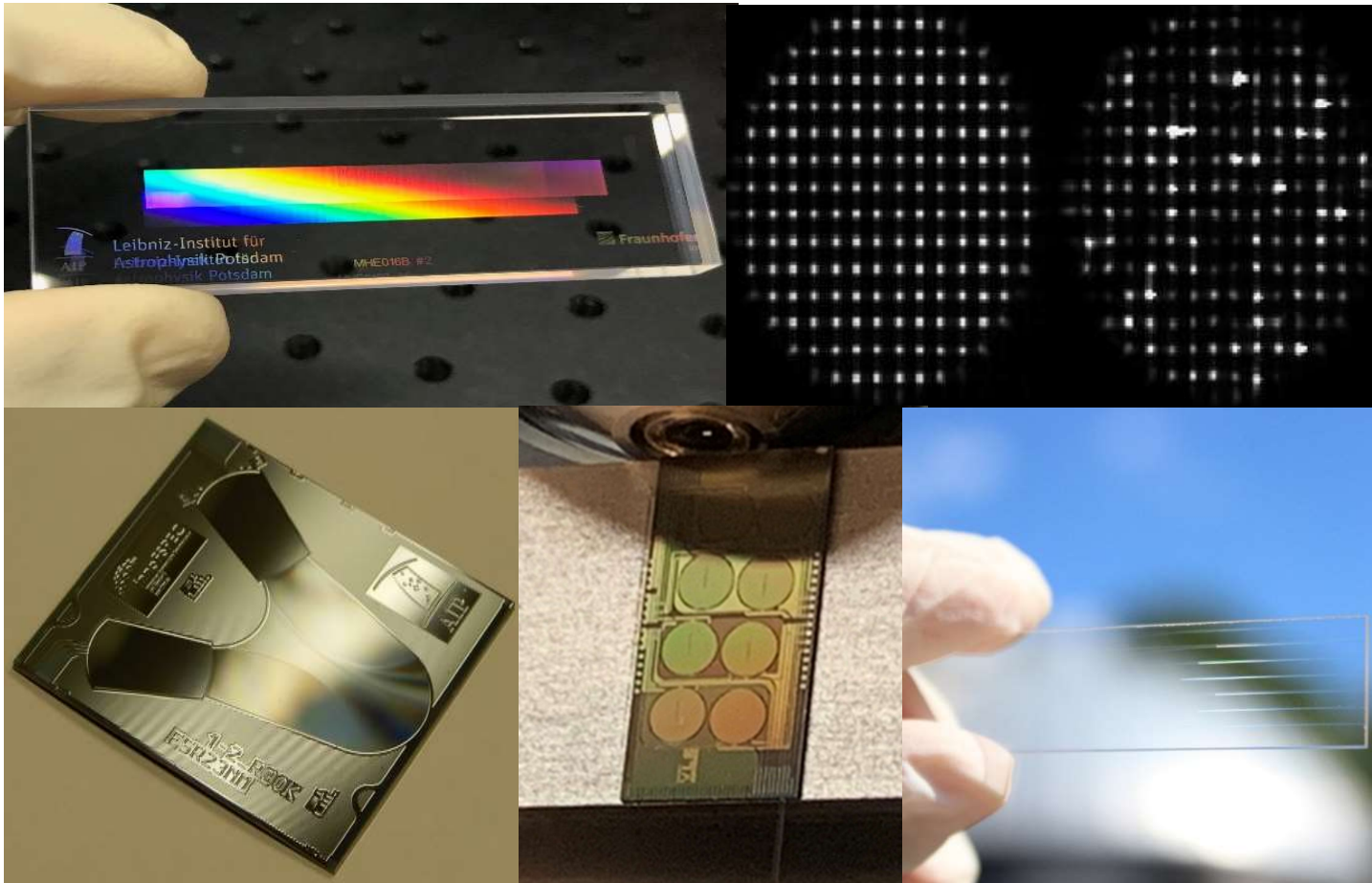


Gravity beam combiner

- 1) Stability
- 2) 100 x size reduction (reduced payload)
- 3) Scalable, faster manufacture
- 4) Less mass and power

Astrophotonics at AIP

- Fiber and chip based photonics for near-infrared astronomy

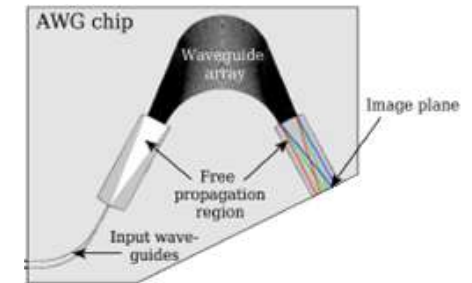
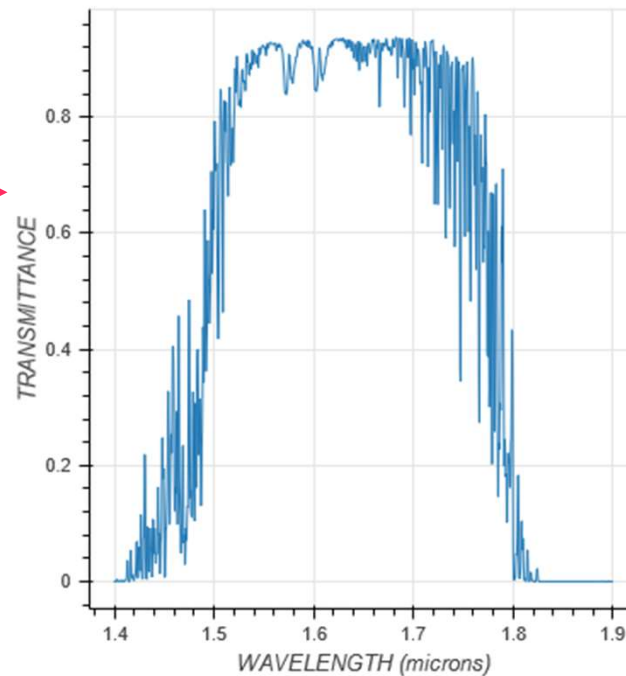
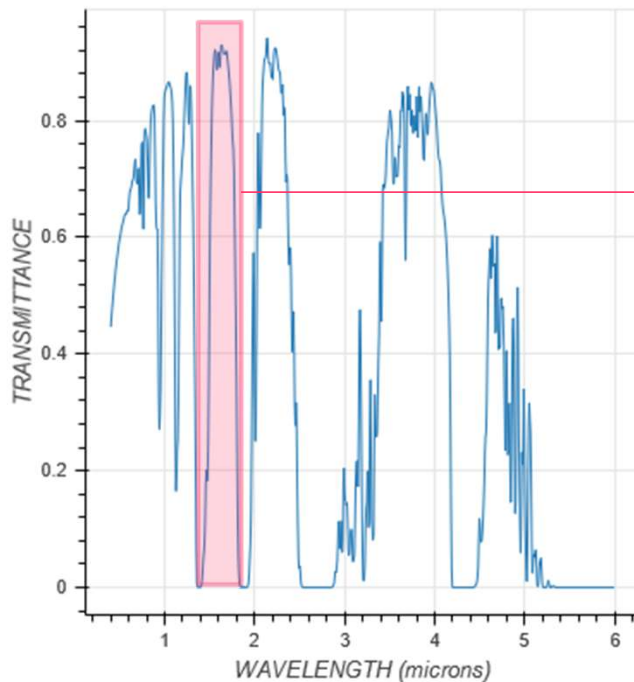


- 1 OH Suppression Filters & Phase Masks
- 2 Adaptive Optics & Photonic Lanterns
- 3 Arrayed Waveguide Gratings
- 4 Frequency Combs
- 5 Pupil Remappers & Beam Combiners

Arrayed Waveguide Gratings

- AWG tailor made for astronomical instrumentation

- 1) Tuned for a broad spectral range – H-Band
- 2) Low insertion loss
- 3) High resolving power



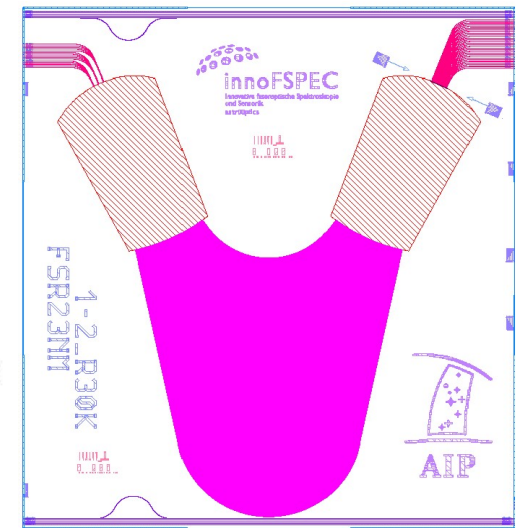
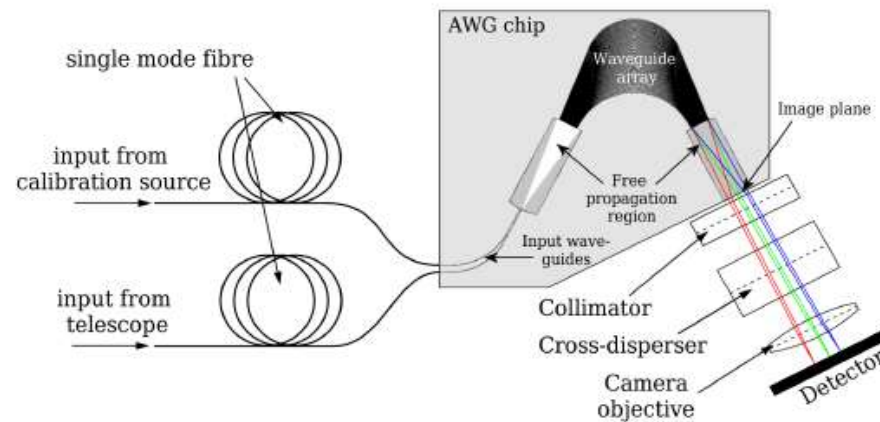
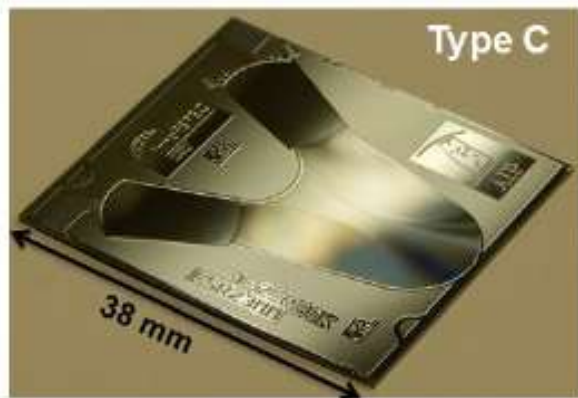
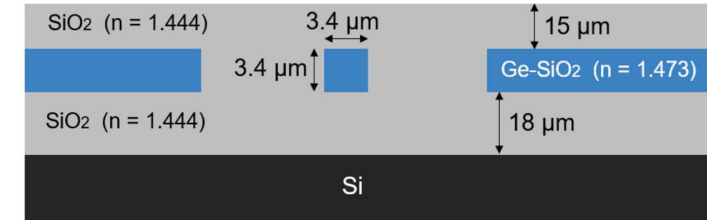
$$\alpha_{\text{losses,dB}} \sim 0.2\text{dB}$$

$$R_{\text{theoretic}} \sim 65000$$

$$R_{\text{measured}} > 15000$$

Integrated Photonic Spectrograph

- Fabrication process
 - 1) UV-Photolithography
 - 2) Atmospheric pressure chemical vapor deposition (APCVD)
 - 3) SiO₂ substrate – Silica on silicon platform
- Best performing chips selected

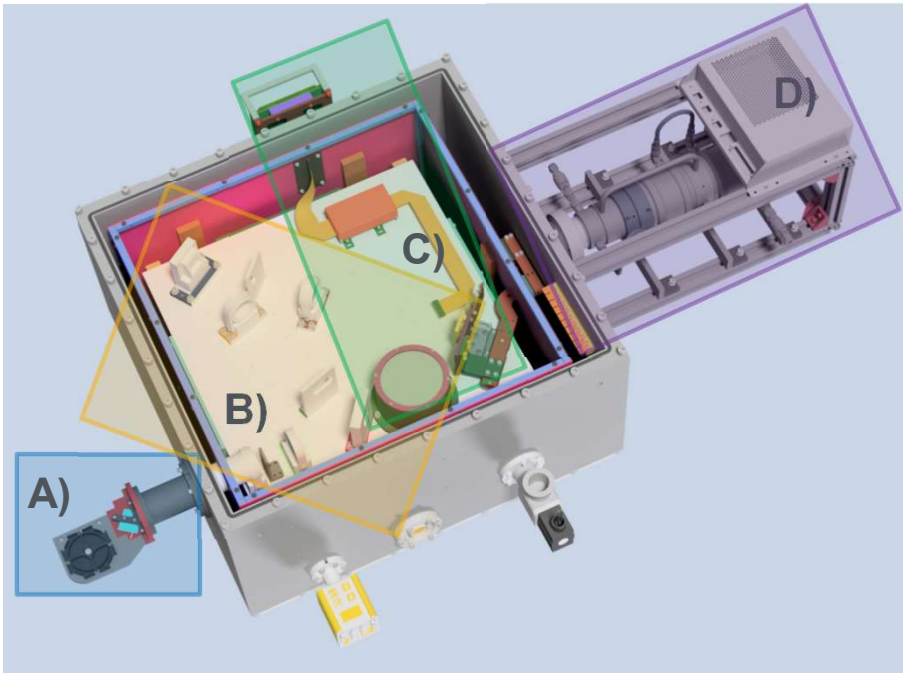


Layout 1-1-2

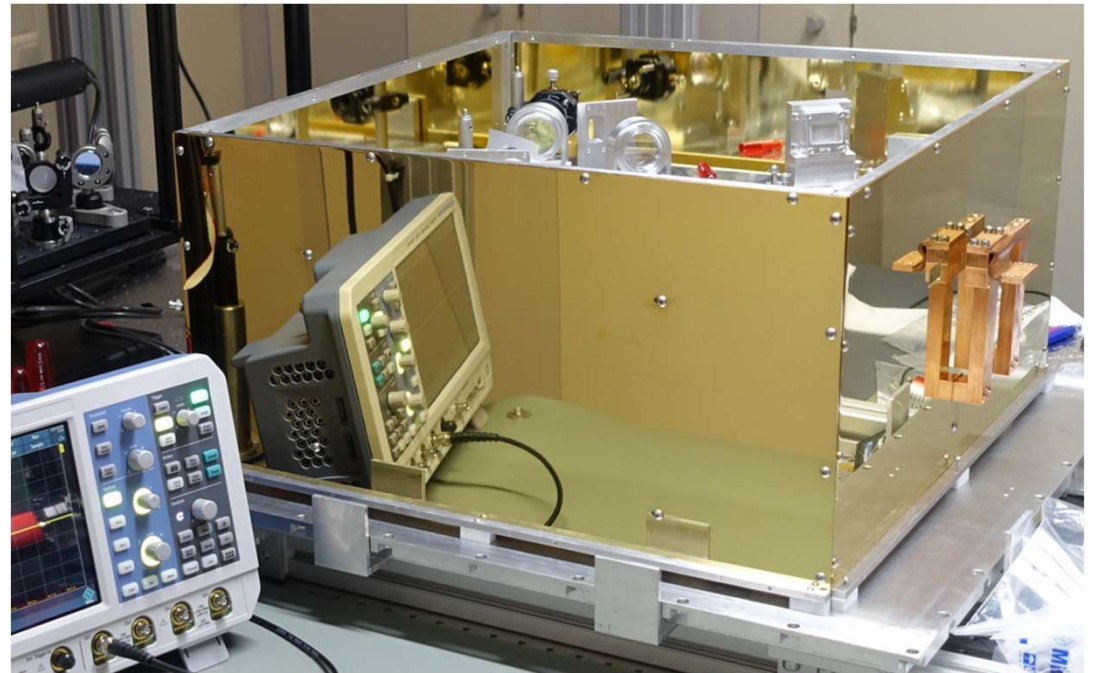
- Integrated photonic spectrograph (IPS)

The Potsdam Arrayed Waveguide Spectrograph

- Design of the cryostat



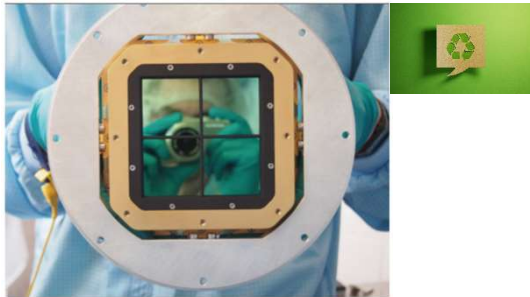
CAD of PAWS



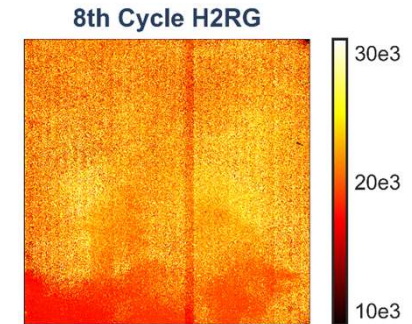
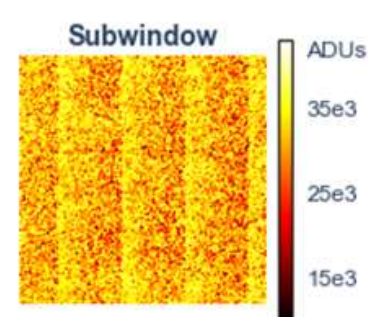
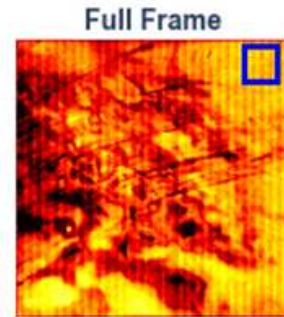
Radiation shield with optical system inside

The Detectors

- The PAWS detector(s)

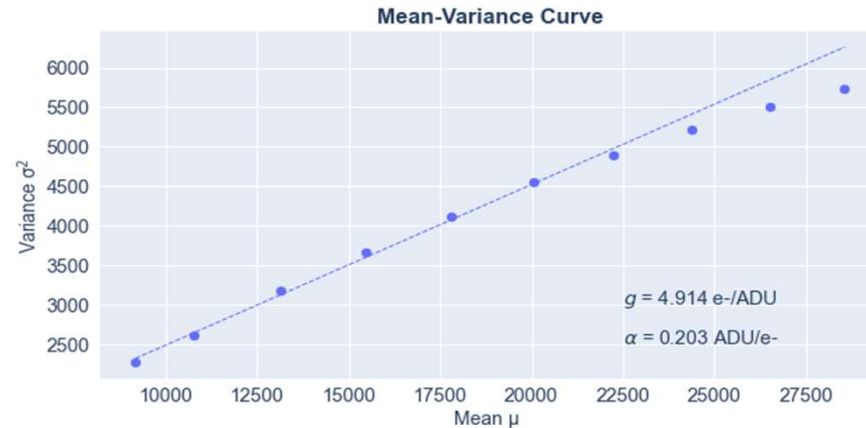


PANIC: 4 H2RG mosaic



Operational temperatures

9th cycle on-going

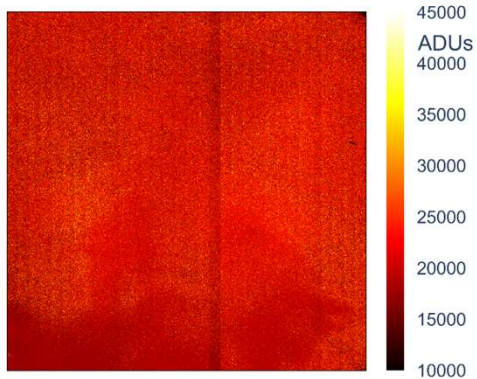


PTC

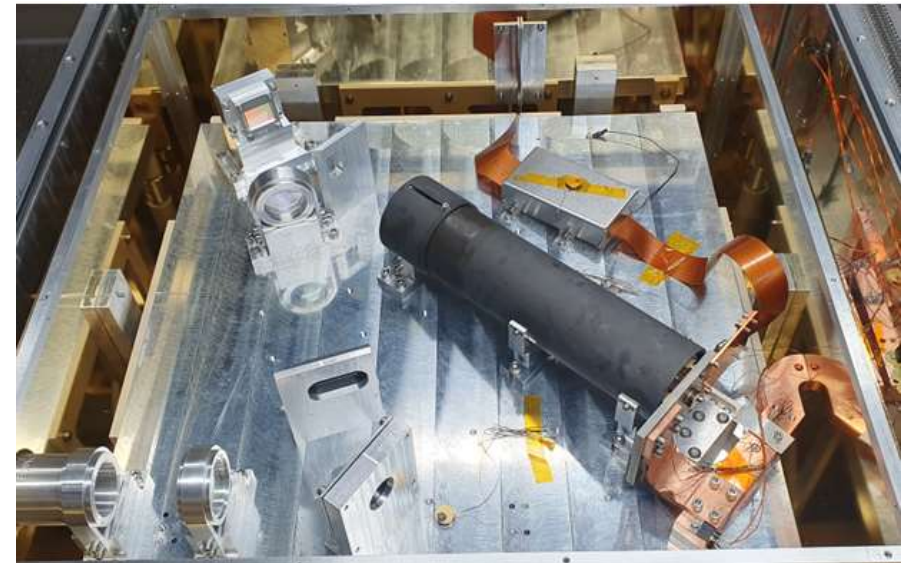
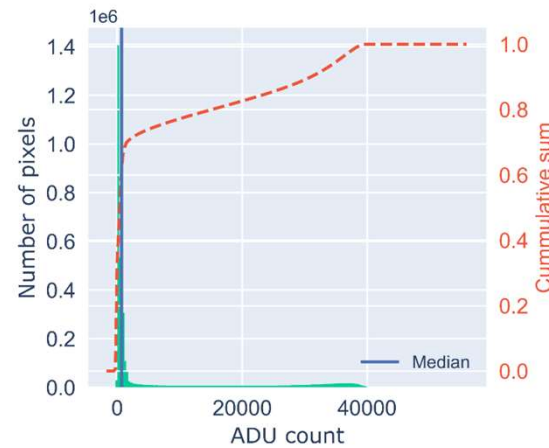
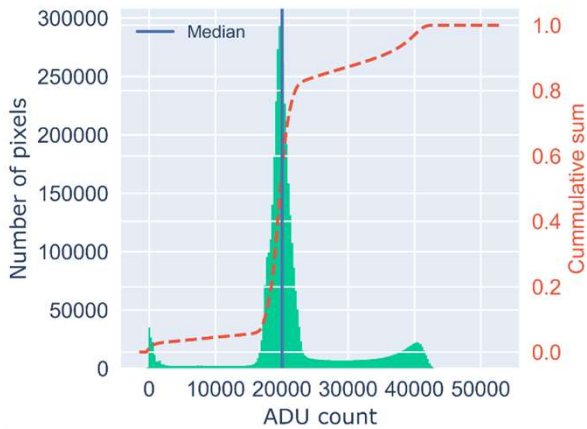
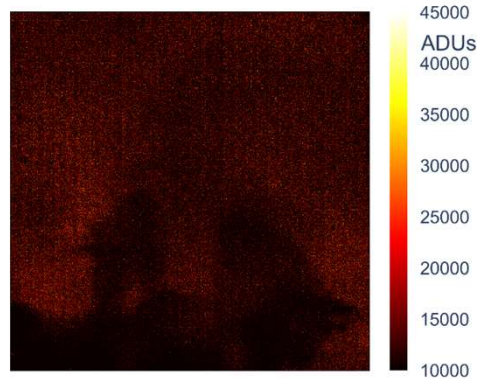
Improvement in Current Cycle

- CDS frames with minimal integration time = 2.7 seconds

8th Cycle



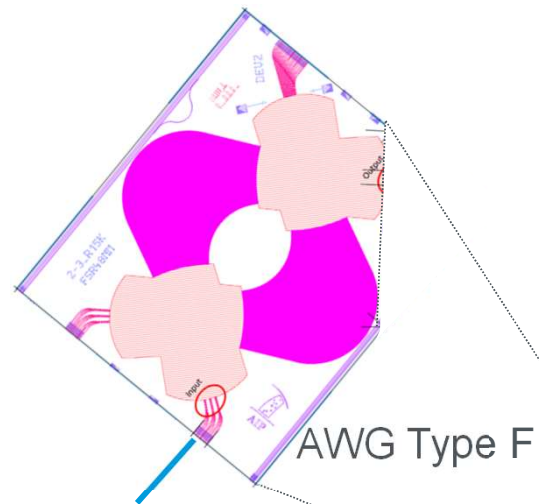
9th Cycle



Baffled detector

First Lab Light

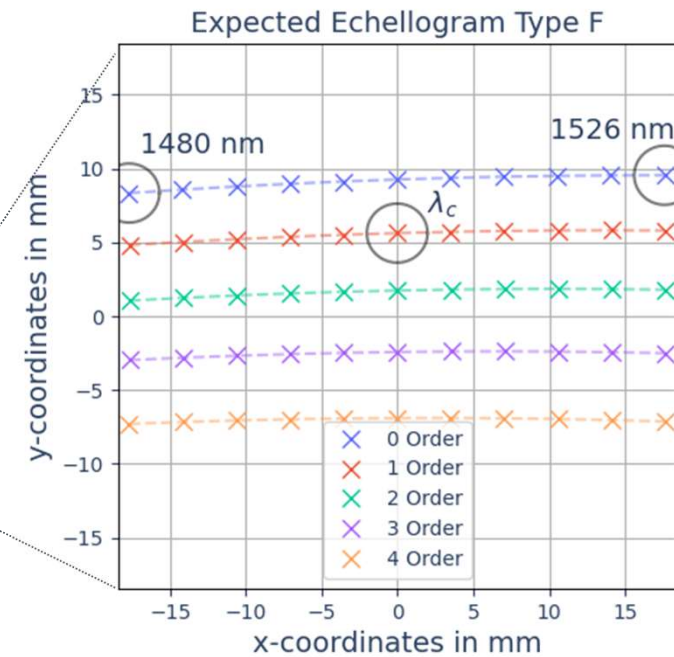
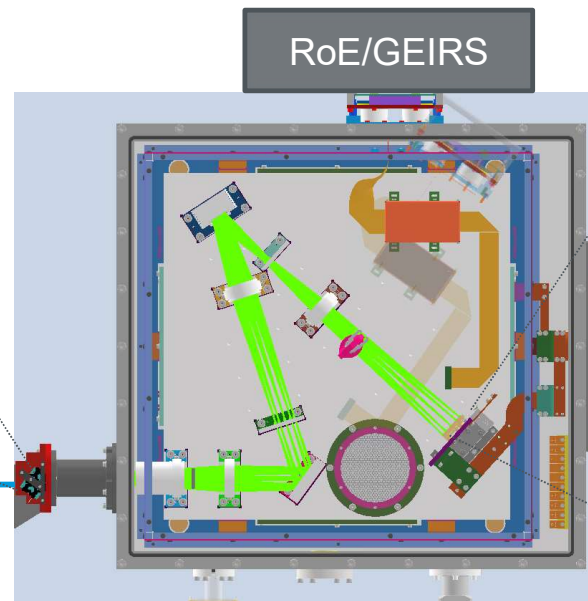
- Experimental Set-up



Tuneable Laser
1400 – 1680 nm
Super-K
1400 – 1800 nm

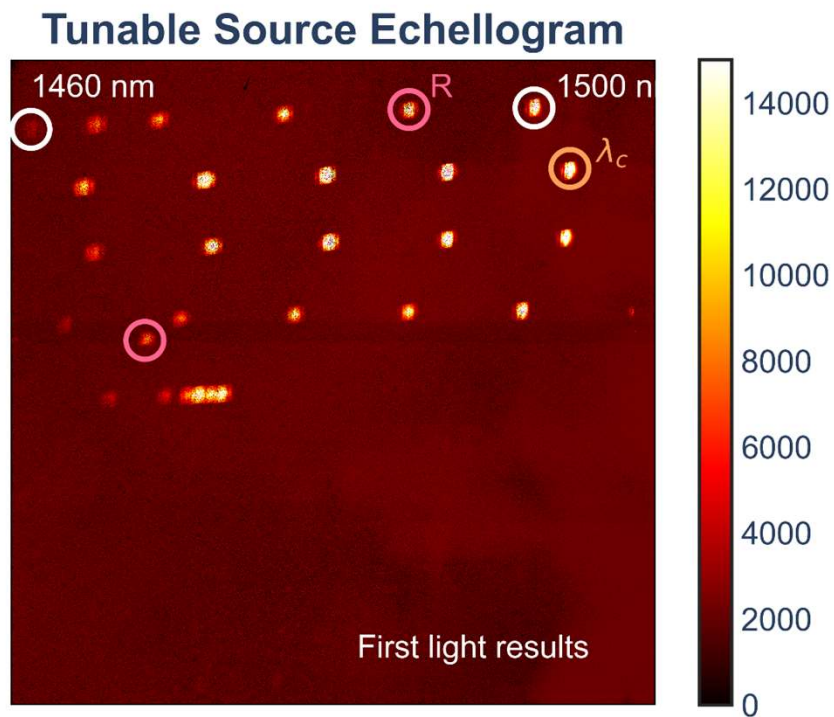
Fibre

Luminos
6 Axis Positioner

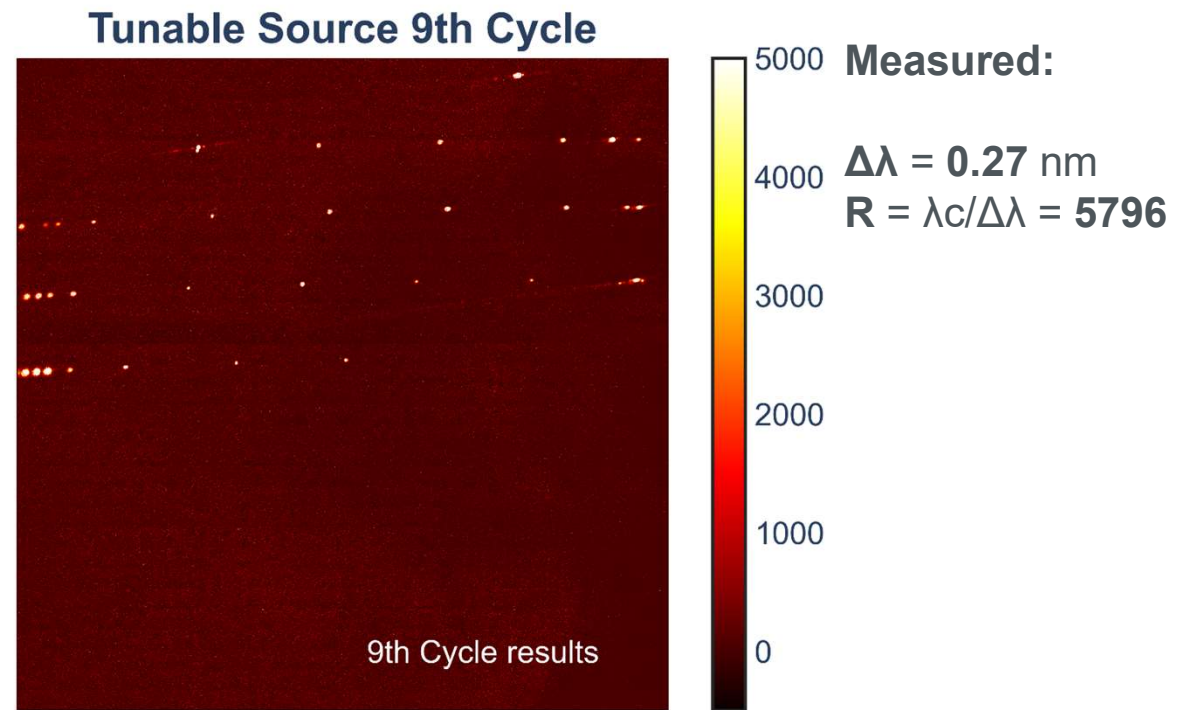


First Lab Light

- Tunable Laser and Super-K results

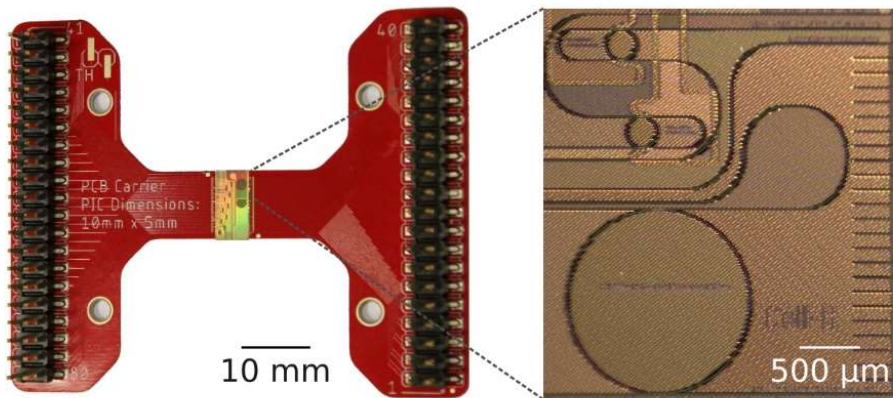
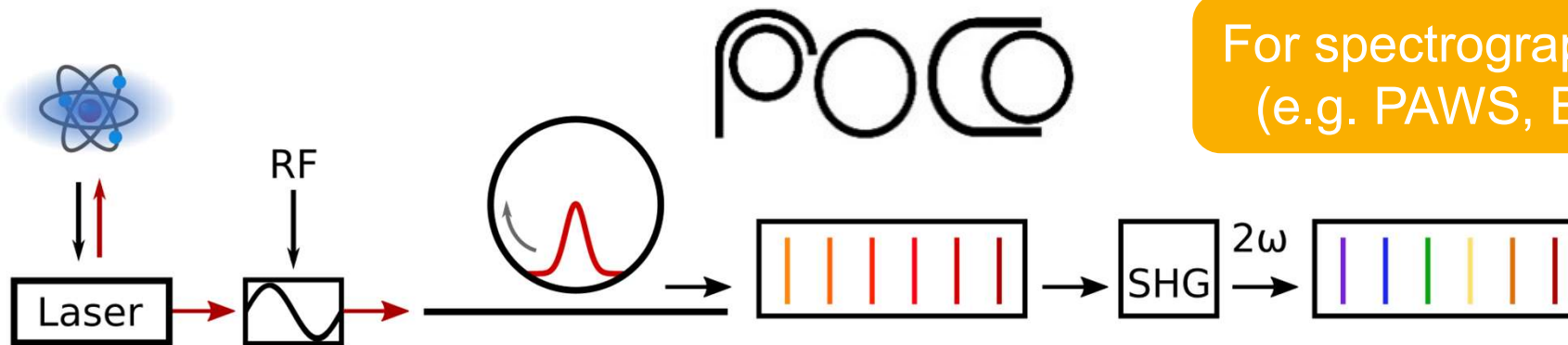


8th cycle rms spot size = 55.1 pixels

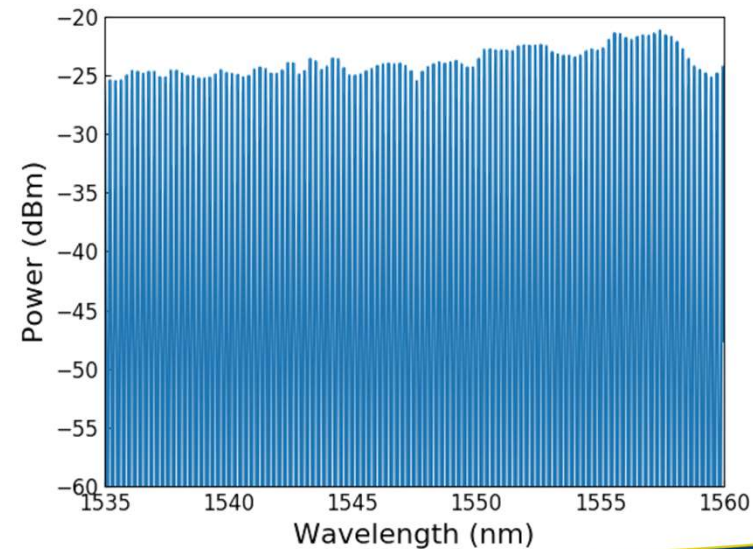


9th cycle rms spot size = **17.43 pixels**
Improvement of factor **3.16**

Chip-based Frequency Comb (Astrocomb)

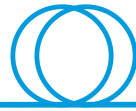


Silicon-Nitride Chip (SiN-Chip) with micro-ring resonator



POCO and PAWS – PICs for Astronomy

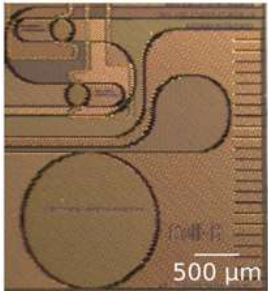
POCO



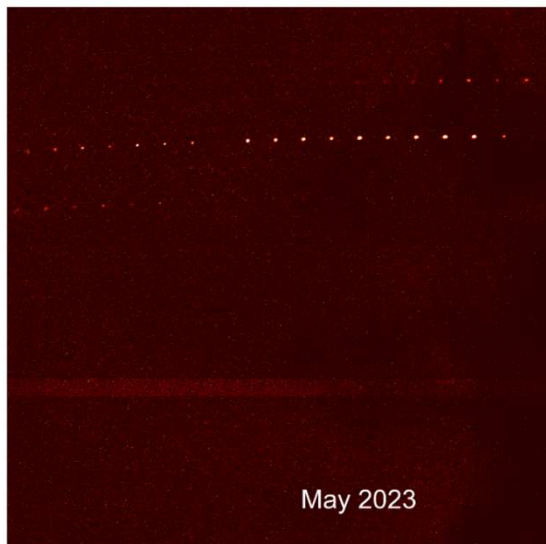
Optical Fibre



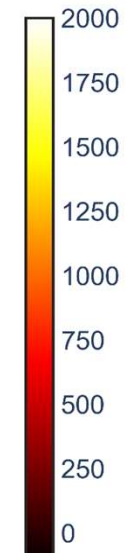
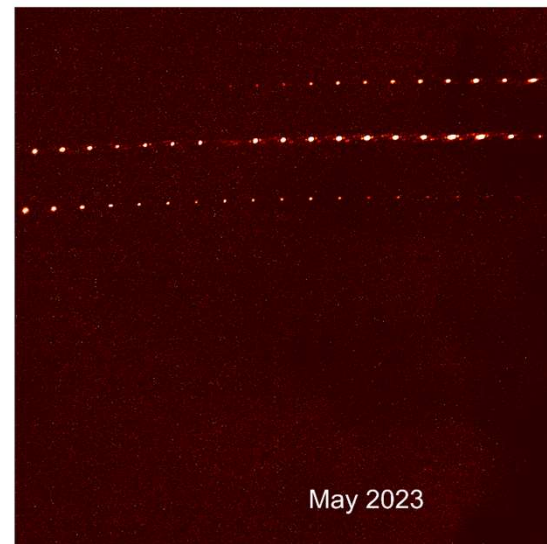
PAWS



POCO-PAWS Measurement

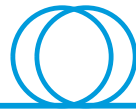
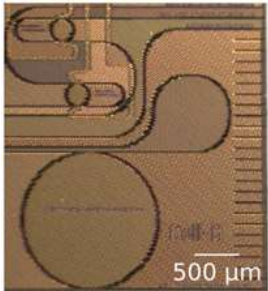


POCO-PAWS Measurement



POCO and PAWS – PICs for Astronomy

POCO



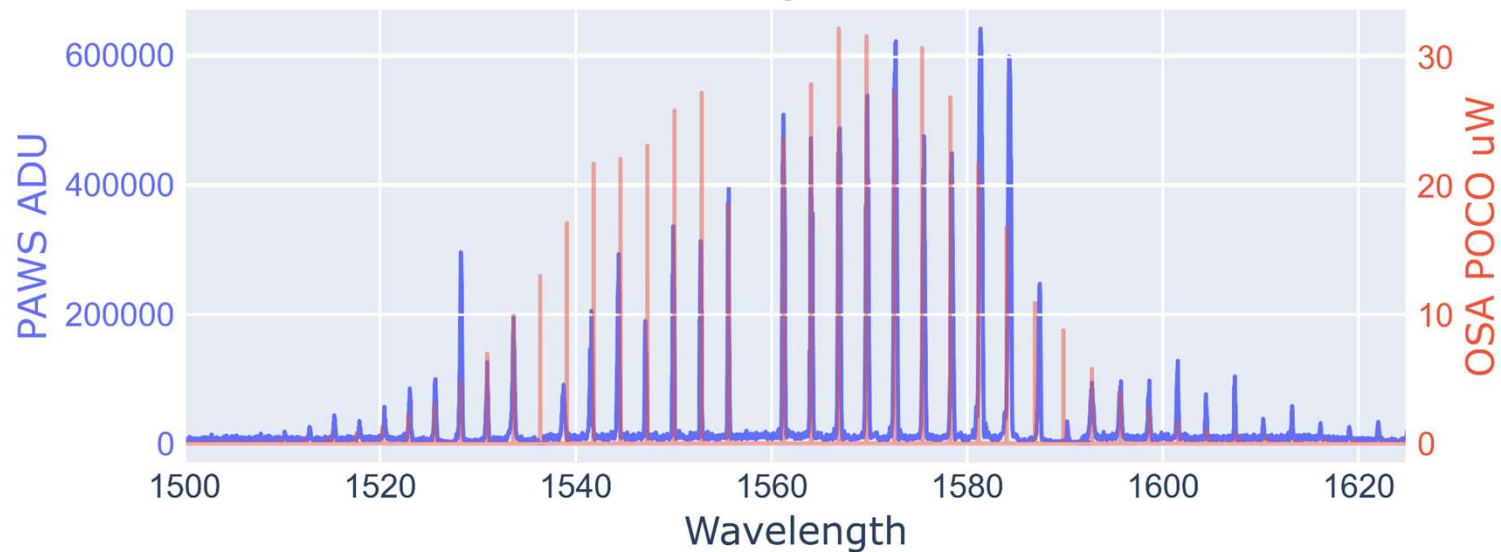
Optical Fibre



PAWS



Measured POCO spectrum with PAWS

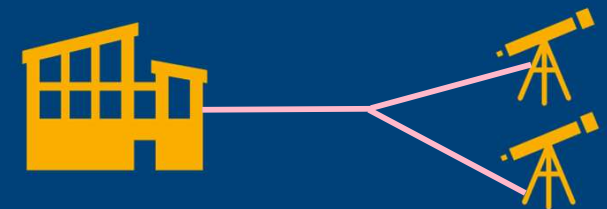


Outlook

- Improvements:
 - 1) Least degraded H2RG
 - 2) Better Focus
- Analysis on AWG performance
 - 1) Test multiple fibres simultaneously
 - 2) Warm AWG limitations
- Test at the Calar Alto Observatory
- Goal: compact cryo-cooled end-to-end system
- Future possibilities of detector integration in PICs for astronomy?

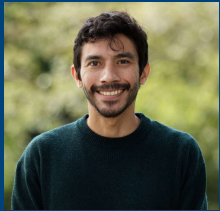
Conclusions

- PICs are used widely in different fields and are constantly improving
- Several benefits for astronomical instrumentation:
 - 1) Stability
 - 2) 100 x size reduction (reduced payload)
 - 3) Scalable, faster manufacture
 - 4) Less mass and power
- Development of Astrophotonic devices on-going:
- Development and first tests of PAWS
- Calibration of PAWS with POCO
- Arrayed waveguide gratings
- Frequency combs
- Interfacility distributed calibration system!
- Goal: compact cryo-cooled end-to-end system



Thanks for listening

PAWS



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Stoll



Stella
Vjesnica



Dr. Alan
Günther

POCO



Daniel
Bondenmüller



Haydar Altuğ
Yıldırım



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Madhav

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